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Automated moment tensor inversionJiří Vackář $^{a 1,2}$ Jan Burjánek 1 Jiří Zahradník 2

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Abstract	Covariance matrix	Synthetic test 1
We are developing a new, fully automated tool for full-waveform moment ten- sor (MT) inversion. It includes automated data retrieval and data selection according to presence of various instrumental disturbances. Data covariance	Inverse problem with no a priory information [Tarantola, 2005]:	$rightarrow$ pure synthetic data — white noise — data + noise C_D inverted R^{1e-21} Z N R^{1e-21} Z N

matrix genereted from before-event noise serves as an automated frequency filter and station weighting according S/N ratio. The software is programmed as much versatile as possible in order to be applicable in other regions and for events ranging from local to regional. It shares some similarities with the broadly used ISOLA software in terms of the inversion methods and in-put/output file structures, but most codes have been re-written from the scratch for maximum computational efficiency. Opposed to ISOLA, whose advantage is in a friendly manual processing of individual events using Matlab GUI, the new codes are intended rather for (i) massive automated application on large sets of earthquakes and/or (ii) near real-time applications.

Goals of the project

- New software for moment tensor inversion, especially for:
- (Near) real-time applications
- Processing large datasets of historical data into moment-tensor catalogs

Properties of the software

- Automated: no user interaction necessary
- Robust: check data quality
- Scalable: from local to regional events
- Universal: applicable in different regions
 Fast: efficient enough to be run in near real-time
 Reliable: evaluate the uncertainty and quality of the result

$\widetilde{\mathbf{m}} = \left(\mathbf{G}^T \mathbf{C}_D^{-1} \mathbf{G}\right)^{-1} \mathbf{G}^T \mathbf{C}_D^{-1} \mathbf{d}_{obs} \qquad (1)$

- model parameters (result)
- data vector
- forward problem matrix
- data covariance matrix

The matrix is calculated from auto-/cross- covariance of before-event noise. The matrix works as automated frequency filter and station weighting to emphasize the high-SNR data.

We started with the covariance matrix for Gaussian random stationary with zero mean [Tarantola, 2005, Example 5.1].

$$\mathrm{C}_D = egin{pmatrix} c(au_0) & c(au_1) & \cdots & c(au_{n-1}) \ c(au_1) & c(au_0) & \cdots & c(au_{n-2}) \ dots & dots & \ddots & dots \ c(au_{n-1}) & c(au_{n-2}) & \cdots & c(au_0) \end{pmatrix} \;.$$

(2)

(4)

We assume the seismic noise to be Gaussian, zero-mean, and stationary, so the last covariance matrix is valid for one component of a station. With assimption of ergodicity, the covariance may be avaluated as correlation (auto-correlation in this case). For discrete time series, the correlation has form

$$f^{fg}(\tau) = (f \star g)[\tau] \stackrel{\text{def}}{=} \frac{1}{2N+1} \sum_{m=-N}^{N} f^*[m] g[m+\tau] .$$
 (3)

The covariance matrix for more stations (here only two components per station for brevity) is then

$$\mathcal{C}_D = egin{pmatrix} \mathrm{C}_{st1}^{NN} \ \mathrm{C}_{st1}^{NE} & 0 & 0 \ \mathrm{C}_{st1}^{EN} \ \mathrm{C}_{st1}^{EE} & 0 & 0 \ 0 & 0 \ \mathrm{C}_{st1}^{EN} \ \mathrm{C}_{st1}^{EE} \ \mathrm{C}_{st2}^{NN} \ \mathrm{C}_{st2}^{NE} \ \mathrm{C}_{st2}^{NE} \ 0 & 0 \ \mathrm{C}_{st2}^{EN} \ \mathrm{C}_{st2}^{EE} \ \end{array}
ight
angle \ ,$$

Where blocks on the diagonal C_{stZ}^{XX} are given by eq. 2 and non-diagonal blocks C_{stZ}^{XY} are similar, only the correlation $c(\tau)$ is replaced by cross-covariance $c^{XY}(\tau)$ between components. Behind the zeroes (zero block) is an assumption, that the seismic noise is not correlated between the seismic stations.

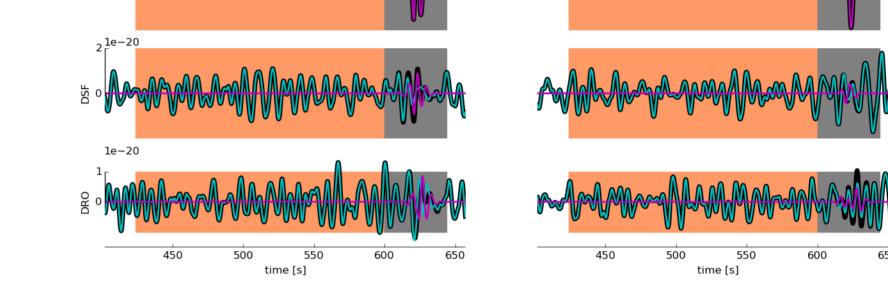


Figure 5: We add strong white noise to 6 of 10 stations. Time windows of the inversion and for calculating the covariance matrix are highlighted.

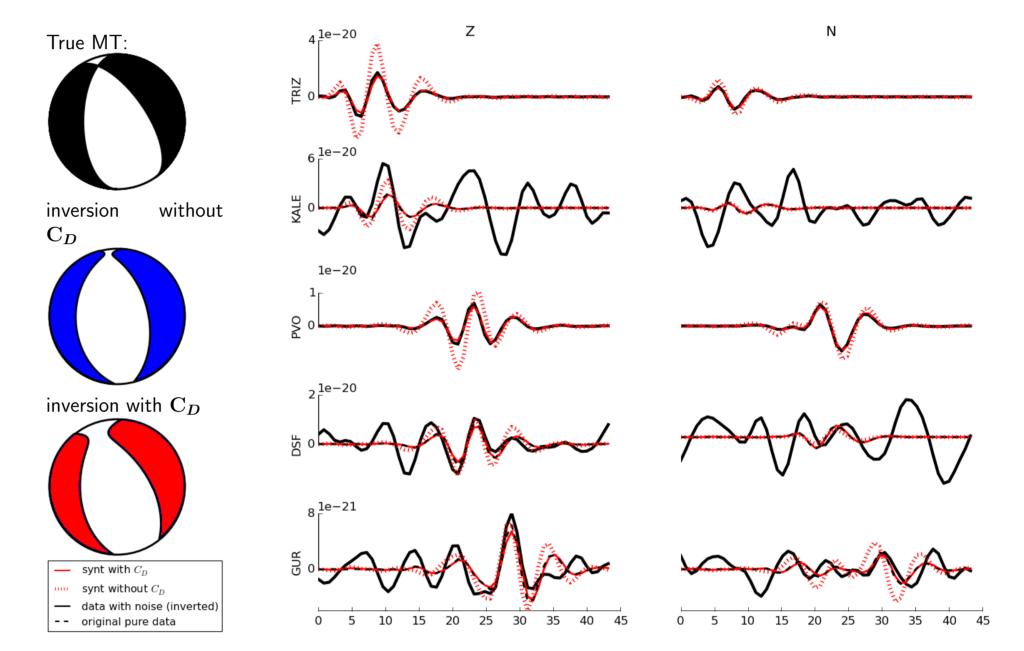


Figure 6: The covariance matrix, which downweight noisy stations, improves the result in terms of waveform fit and agreement of strike/dip/rake with the original MT. The CLVD % seems to be less stable part of the result.

Technical solution

Programmed as *Python* module, using *ObsPy* framework:

ISOLA-ObsPy documentation »		next modules index
Table Of Contents	ISOLA-ObsPy documenta	tion
ISOLA-ObsPy documentation • Function summary • Requirements • Download • Installation	ISOLA-ObsPy is an open-source module for Python for solution of seismic source inverse problem. It uses the point source approximation and describe the source in form of moment tensor. Copyright: Jiří Vackář	
 Examples Indices and tables 	Version: developer's snapshot 2015-06-11	
Next topic	License: GNU Lesser General Public License, V	ersion 3 (http://www.gnu.org/copyleft/lesser.html)
class_isola module	Contents:	
This Page	class isola module	
Show Source	-	
Quick search	Function summary	
	ISOLA([location_unc, depth_unc, time_unc,])	Class for moment tensor inversion.
Go	Top 1 1(of nowline])	Write text into log file
	ISOLA.log(s[, newline])	
GC Enter search terms or a module, class or function name.	ISOLA. rog(s[, newinte]) ISOLA. read_crust (Source[, output])	Copy a file with crust model definition to location where code Axitra expects it
Enter search terms or a module,		

Figure 1: Webpage with the documentation

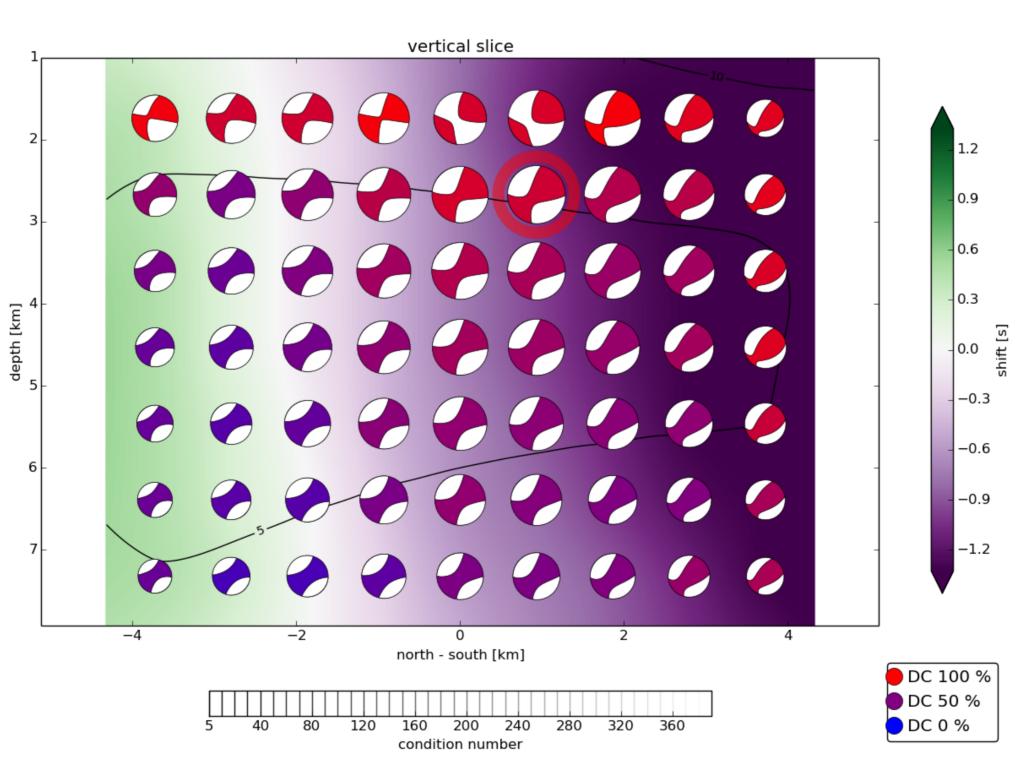
Program schema

- ▶ input: location, approx. magnitude
- ► load network configuration file / SeisComP DB
- load waveforms & instrument responses files / ArcLink
 disturbance detection

Automatically ploted output

BOARD AND MAN gold when and 5 10 15 20 25 30 35 0 5 10 15 20 25 30 35 0 5 10 15 20 25 30 35

Figure 3: Automatically ploted waveform fit



Synthetic test 2

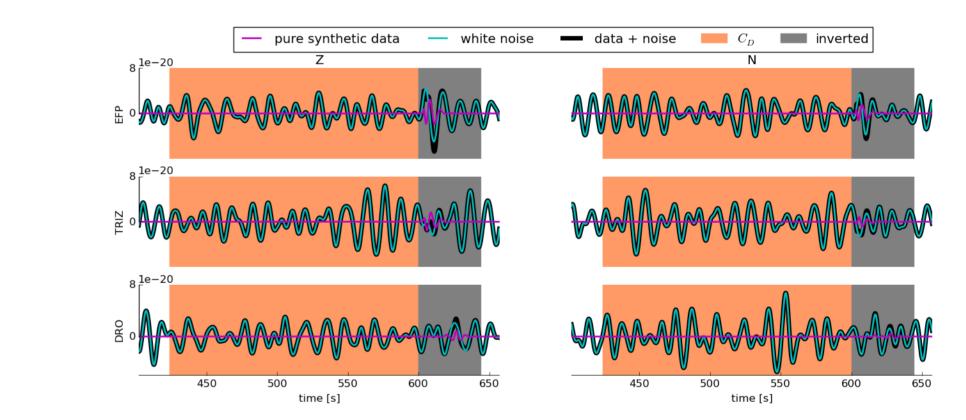


Figure 7: We add strong filtered white noise to all stations. Inverted frequencies are 0.05–0.15 Hz, the noise is filtered by bandpass filter 0.05–0.10 Hz.

True MT version with \mathbf{C}_{L}

- create 3-D space + time grid to seek the centroid
- calculate Green's functions in grid points (Axitra) in parallel
- create covariance matrix
- solve inverse problem in each grid point in parallel
- plot the results

Disturbances detection

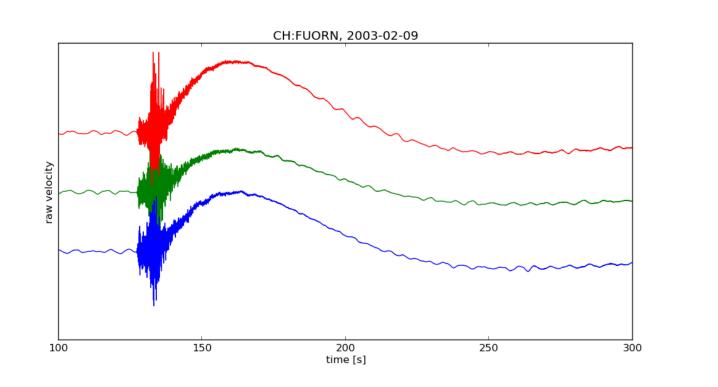


Figure 2: Typical example of long-period disturbance, which is caused by seismometer tilt or instrument malfunction. Such disturbances are detected using *MouseTrap* code [Vackář et al., 2015] and removed from processing Figure 4: North-south slice over grid of solution. The best solution is highlighted. Sizes of beachballs correspond to variance reduction, their colors shows DC %. Color contours in the background shows inverted centroid time, contour lines displays condition number of the inverse problem, which correspond to solution stavbility. Such slices in 4 direction through the point of the best solution are ploted automatically, as well as horizontal slices in each depth.

Acknowledgement

The research was financially supported by SCIEX grant in Switzerland and the following grants in the Czech Republic: SVV 115-9/260096 and GAUK 496213.

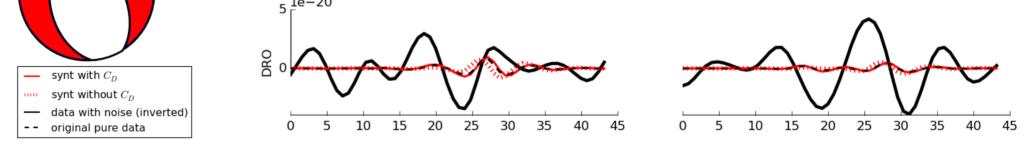


Figure 8: The covariance matrix, which filter out noisy frequencies, comes to the right result. The solution without covariance matrix is completely wrong.

Conclusion

- ► We develop a new software for automated MT inversion
- It is adaptable from local to regional events
- ► We do not need exact location on the input
- Inverse problem formulation with noise covariance matrix
- It improves the results in case of noisy data

References

TARANTOLA, A. (2005), Inverse problem theory and methods for model parameter estimation. siam. VACKÁŘ, J., J. BURJÁNEK, and J. ZAHRADNÍK (2015). Automated detection of long-period disturbances in seismic records; MouseTrap code, *Seismological Research Letters*, 86, 442–450. http://geo.mff.cuni. cz/~vackar/mouse/

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